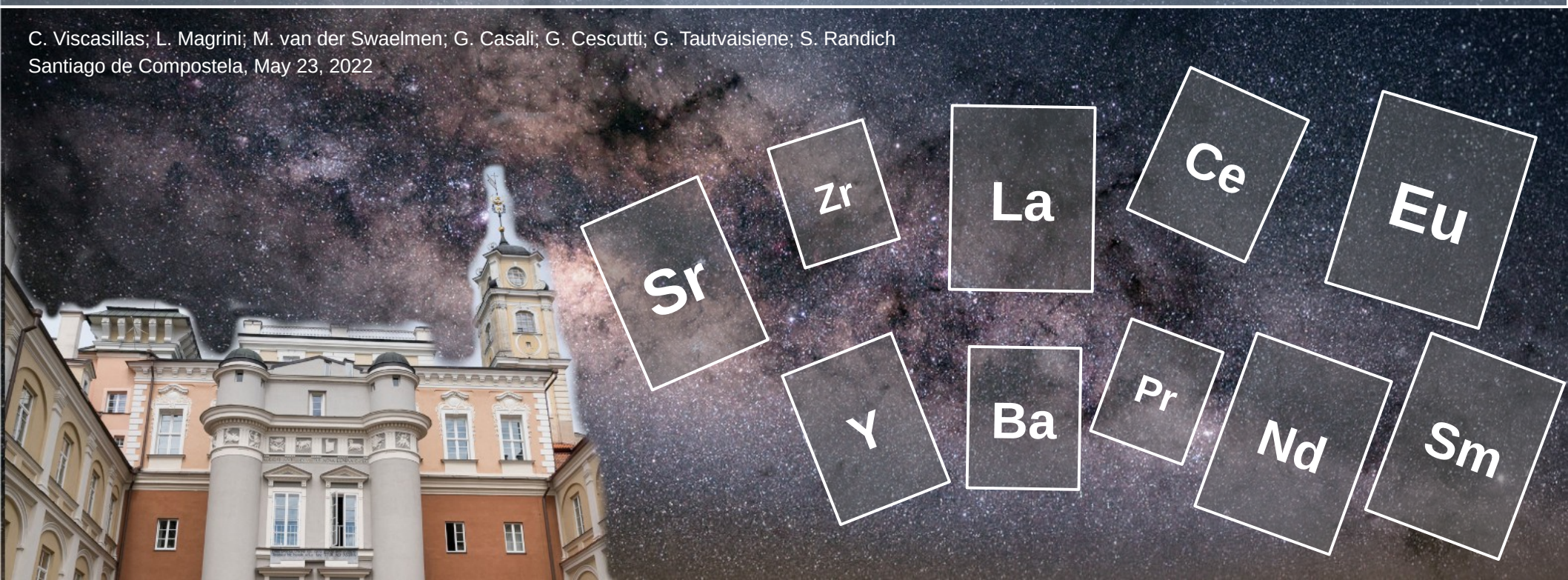


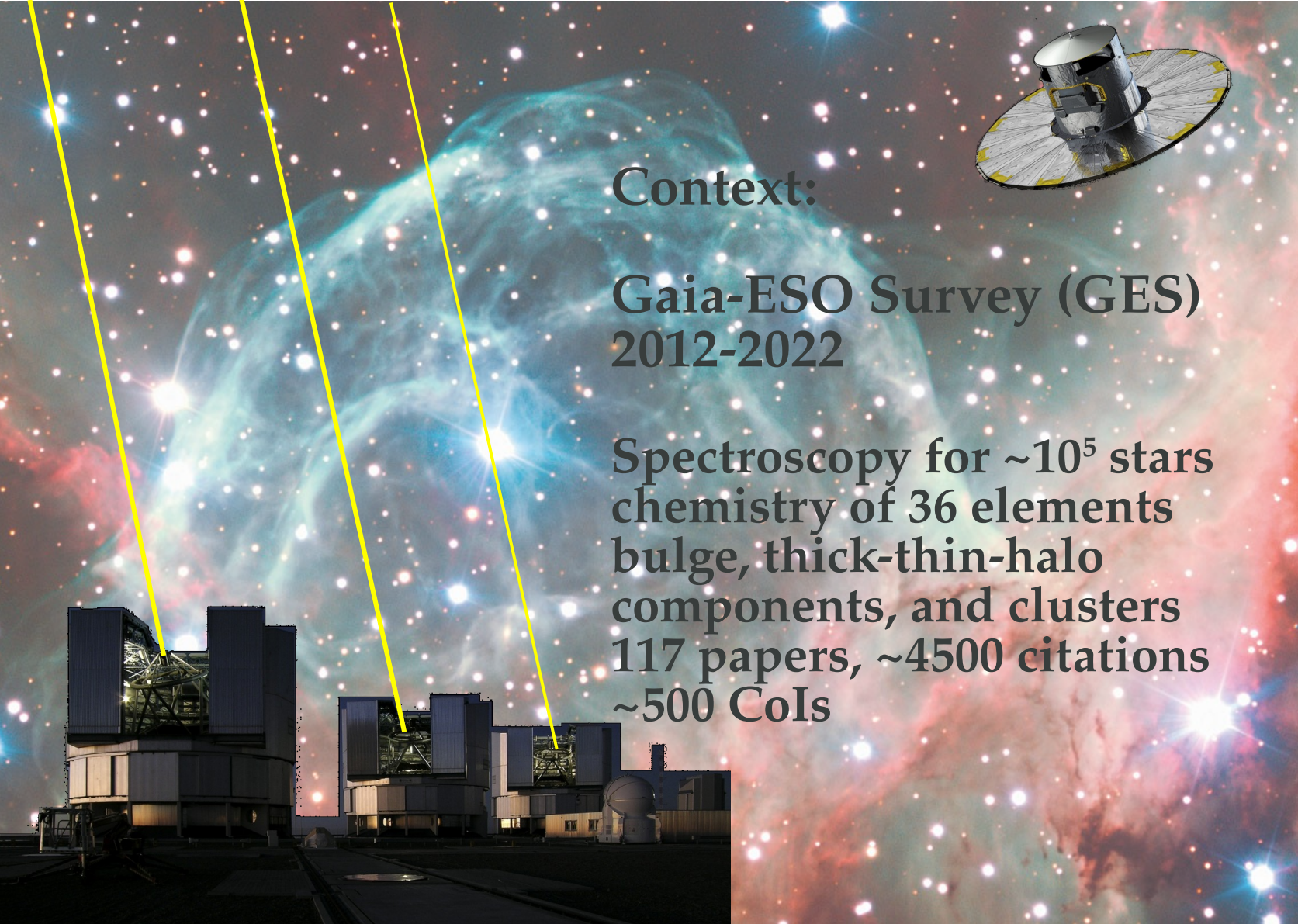
Exploiting the abundances of neutron-capture elements from the latest Gaia-ESO Survey (GES) data release.



Carlos Viscasillas Vázquez (Vilnius University; AO INAF-Arcetri)

C. Viscasillas; L. Magrini; M. van der Swaelmen; G. Casali; G. Cescutti; G. Tautvaišienė; S. Randich
Santiago de Compostela, May 23, 2022

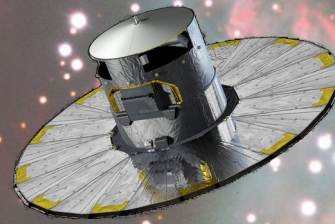




Context:

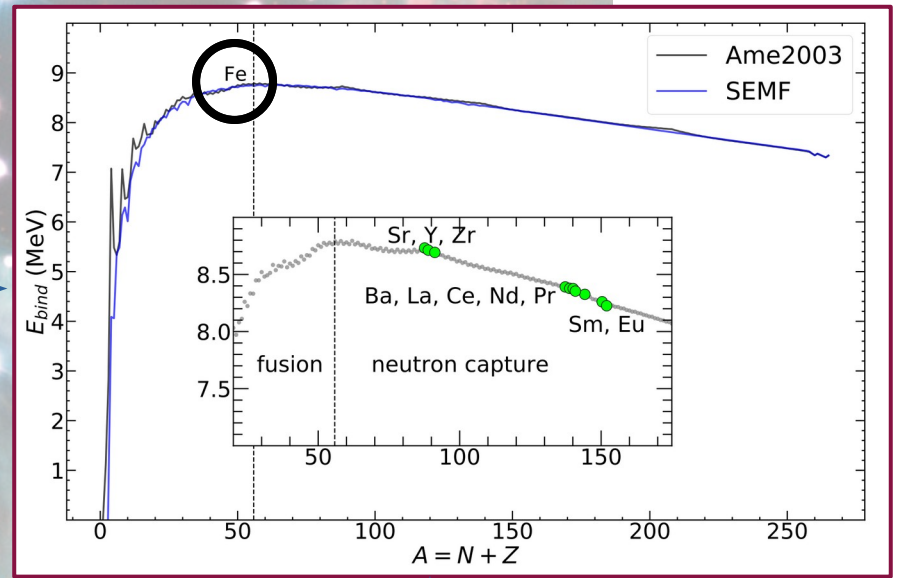
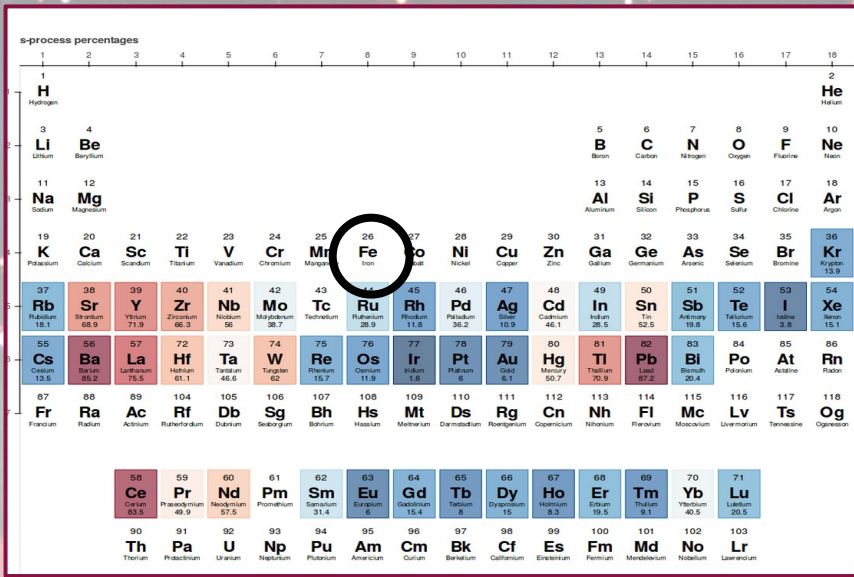
**Gaia-ESO Survey (GES)
2012-2022**

**Spectroscopy for $\sim 10^5$ stars
chemistry of 36 elements
bulge, thick-thin-halo
components, and clusters
117 papers, ~ 4500 citations
 ~ 500 CoIs**

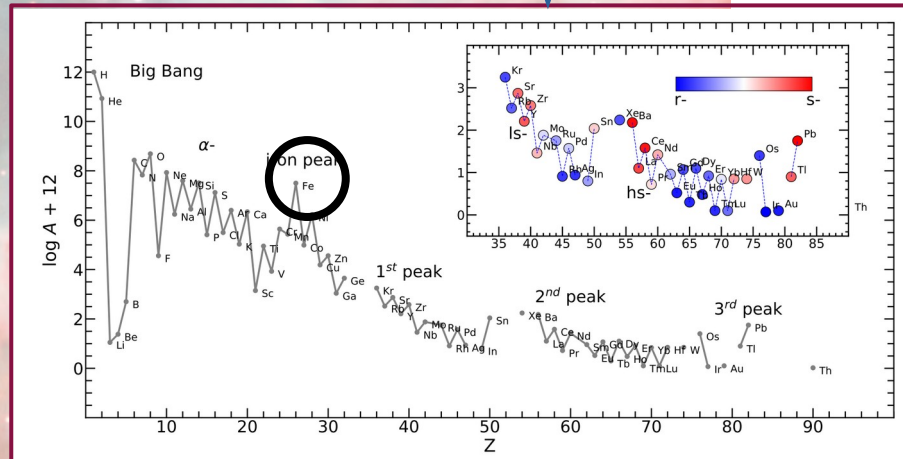


**MW-Gaia Cost
STSM AO
INAF-Arcetri**

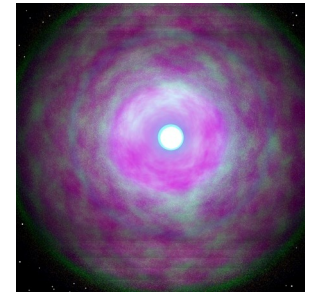
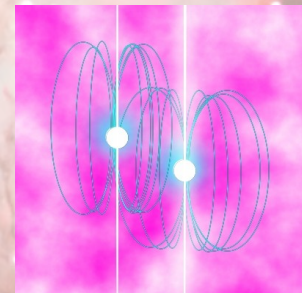
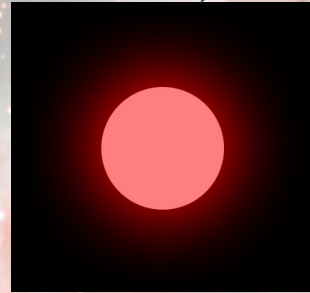
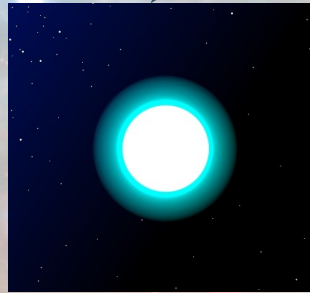
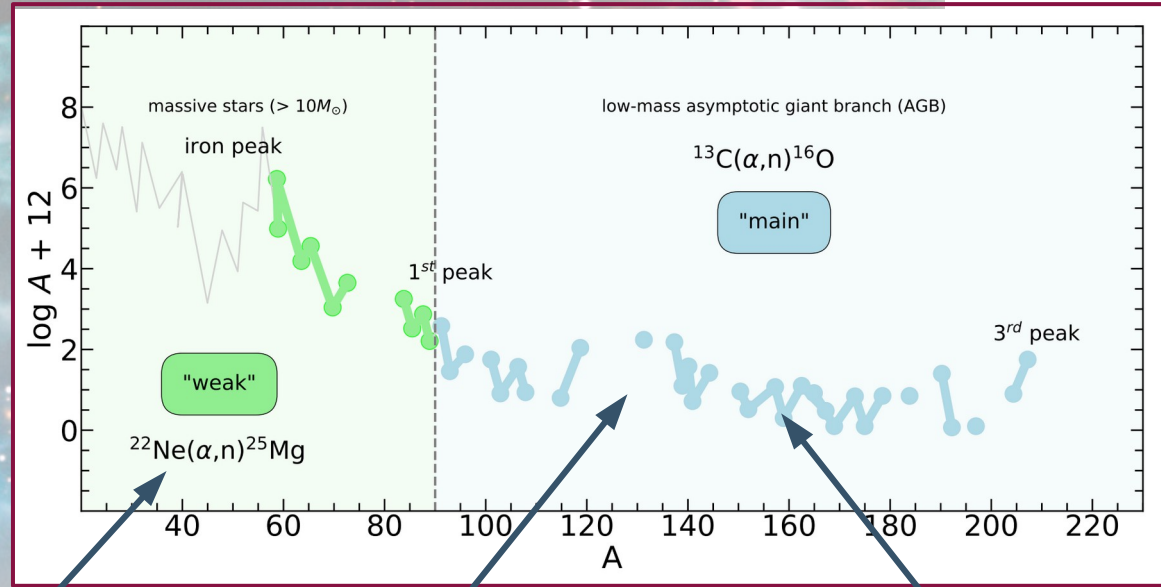




Neutron-capture elements:
special nucleosynthesis
s- and r- process



different origin
and nature
astrophysical sites



applications of *s*-process and problems of *r*-process



Gaia
astrometry-
positions
~ 10^9 stars

+



Gaia
Spectroscopy-
chemistry-
RVs
~ 10^5 stars

+

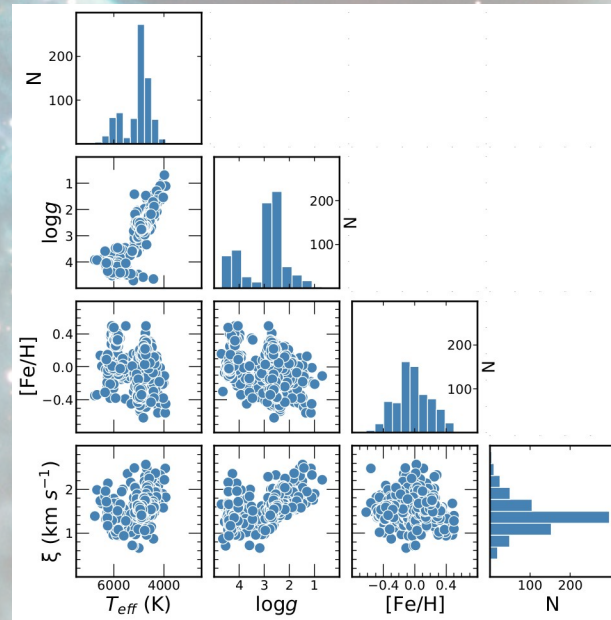
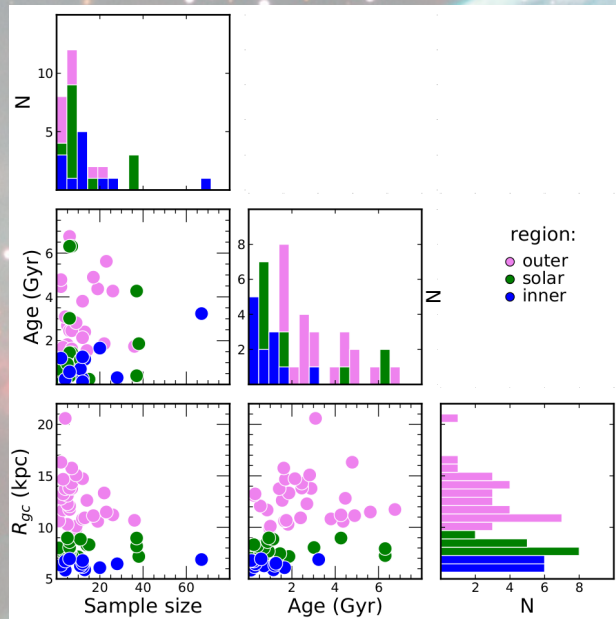


Ages
(missing variable)

GALAH (~600,000 stars)
RAVE (~450,000 stars)
APOGEE (150,000 stars)
Upcoming: MAVIS,
HRMOS, MOONS

The sample:

Viscasillas, Magrini+22 (A&A) GES paper 112

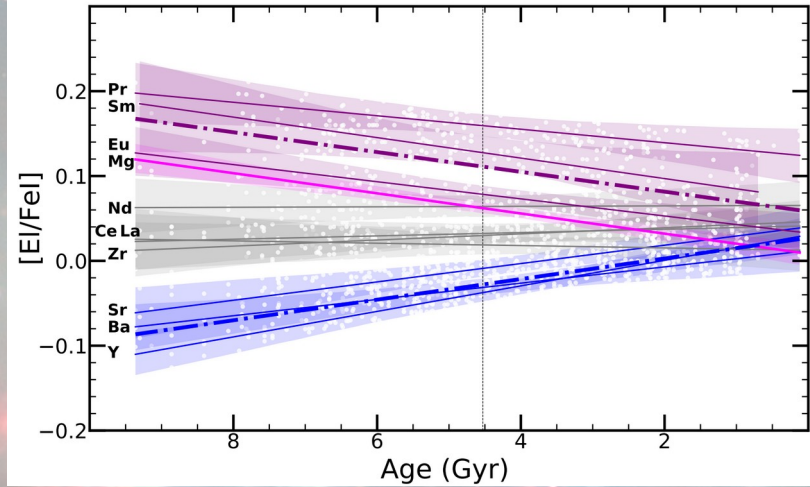


62 open clusters

~750 member stars

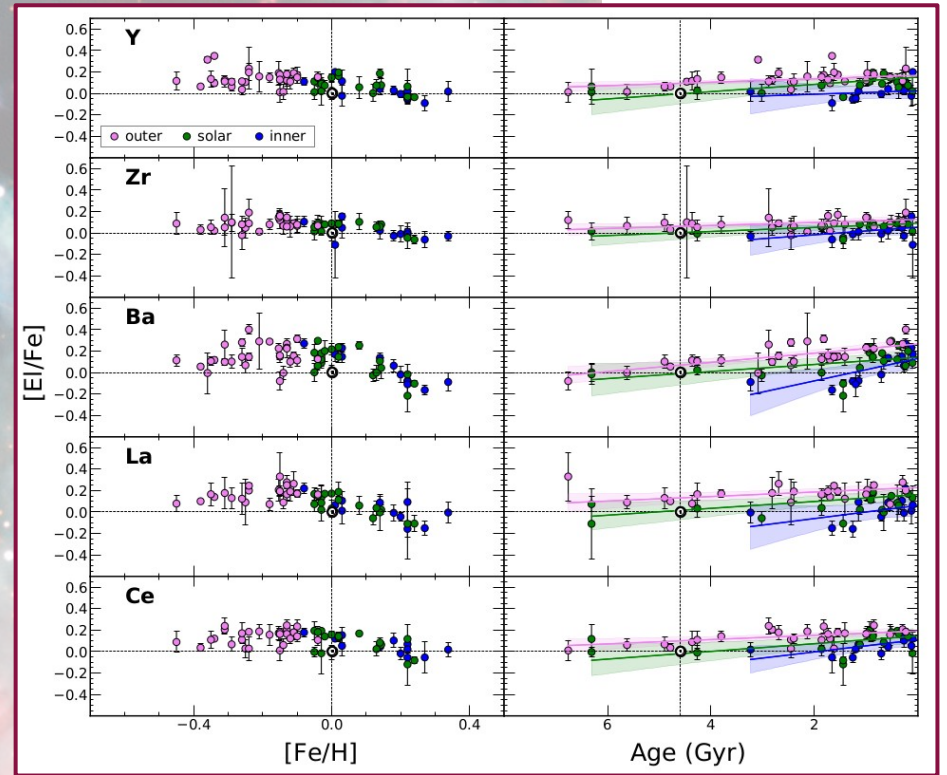
large sample of **open clusters**, most reliable tracers of the chemical evolution of the Galaxy
newest release of the **Gaia-ESO Survey** based on the high resolution spectroscopy of UVES.

Thin disc



Tautvaišienė, Viscasillas, Mikolaitis+21
(A&A)

Open clusters

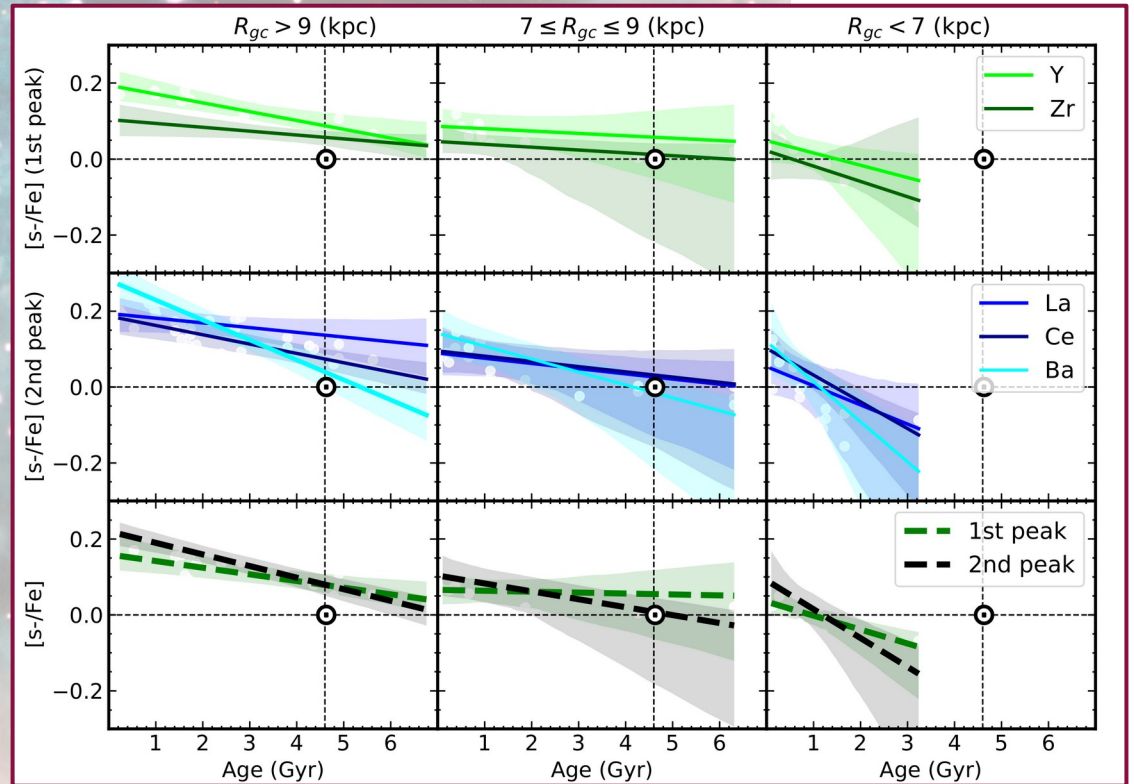


Viscasillas, Magrini+22 (A&A)

Chemical clocks
Sensitivity to age: Y and Ba

Sensitivity to age from another perspective

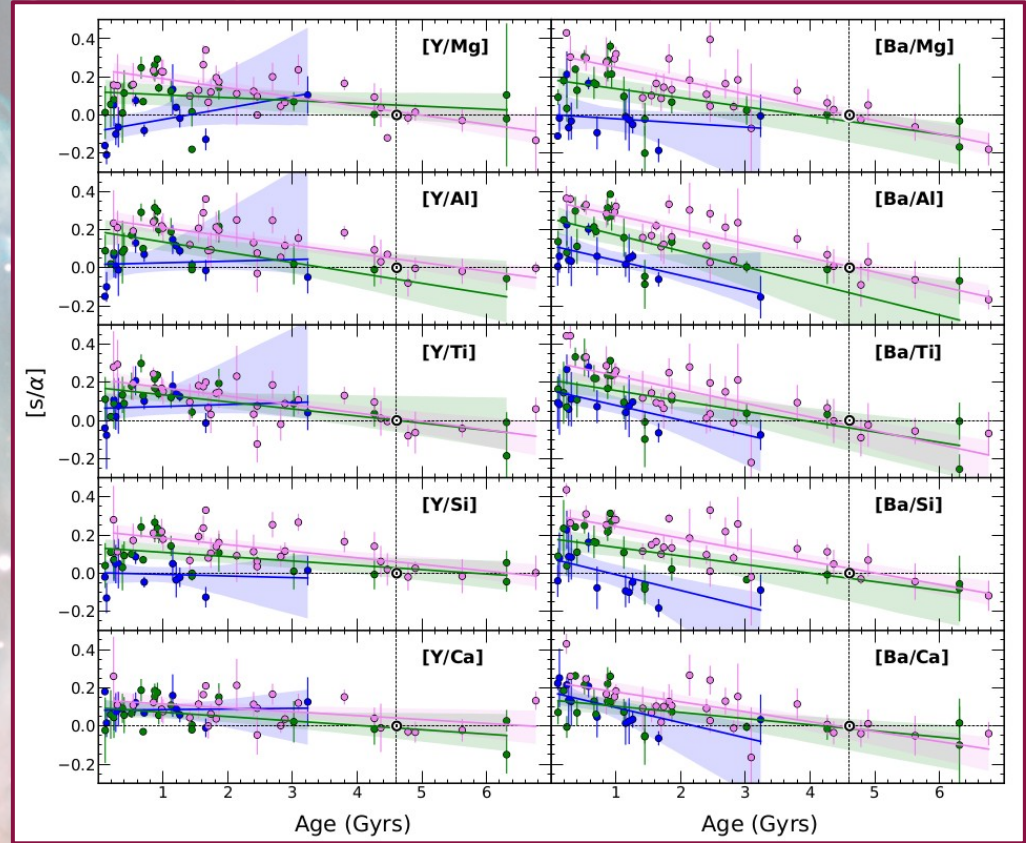
- Metallicity-dependent yields for s-process elements
- Differences between 1st and 2nd peaks



Magrini, Viscasillas+22 (Universe)

Chemical clocks: $[s/\alpha]$

- We provide a set of 40 weighted MLR in 3 variables ($[\text{Fe}/\text{H}]$, R_{gc} , and age).
- We apply 3 tests to our relations:
 - a) in recovering the ages of the open clusters;
 - b) in recovering the ages of member stars;
 - c) to infer ages of field stars in $[\alpha/\text{Fe}]-[\text{Fe}/\text{H}]$ plane.
- We investigated the role of migration in our relations.



Viscasillas, Magrini+22 (A&A)

(kpc)



inner

solar

outer

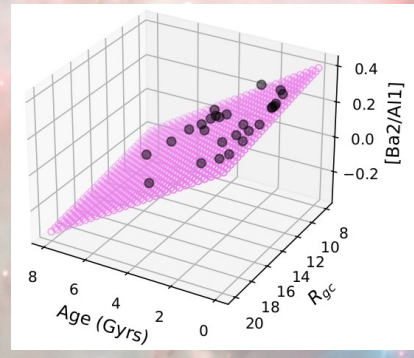
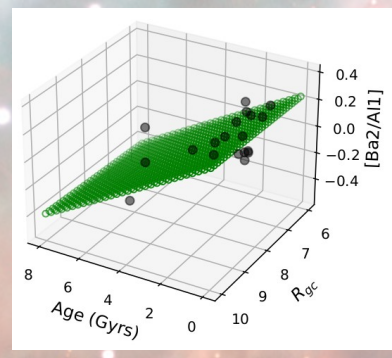
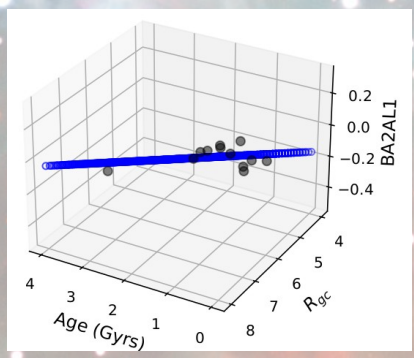
global

$$[s/\alpha] = m_1 \cdot \text{Age} + m_2 \cdot R_{gc} + m_3 \cdot [\text{Fe}/\text{H}] + c$$

$$[s/\alpha] = m_1 \cdot \text{Age} + m_2 \cdot R_{gc} + m_3 \cdot [\text{Fe}/\text{H}] + c$$

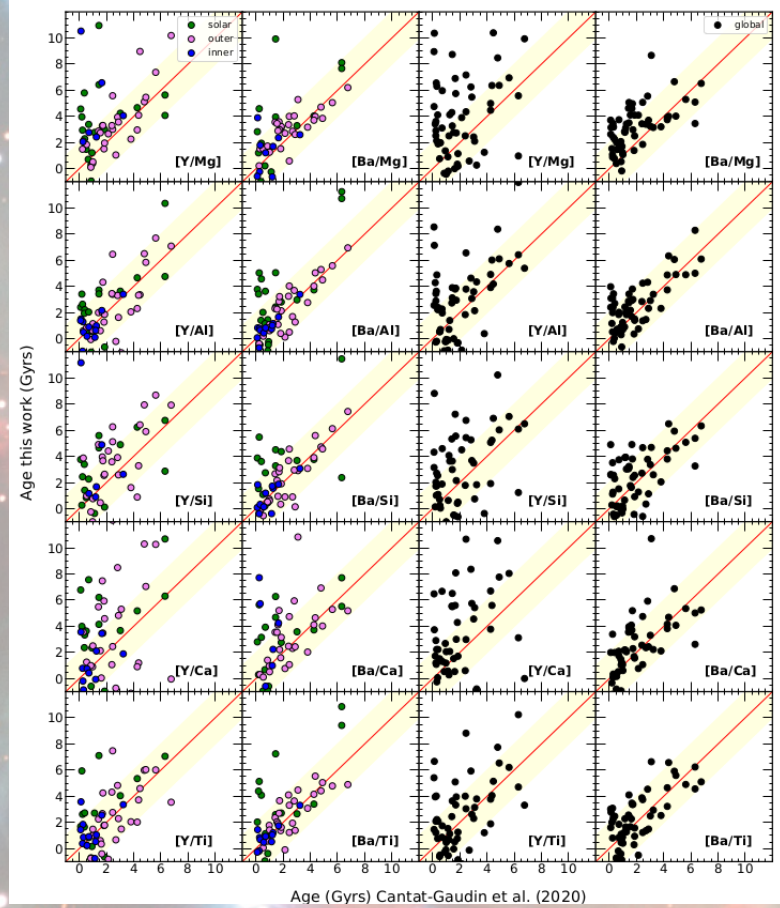
$$[s/\alpha] = m_1 \cdot \text{Age} + m_2 \cdot R_{gc} + m_3 \cdot [\text{Fe}/\text{H}] + c$$

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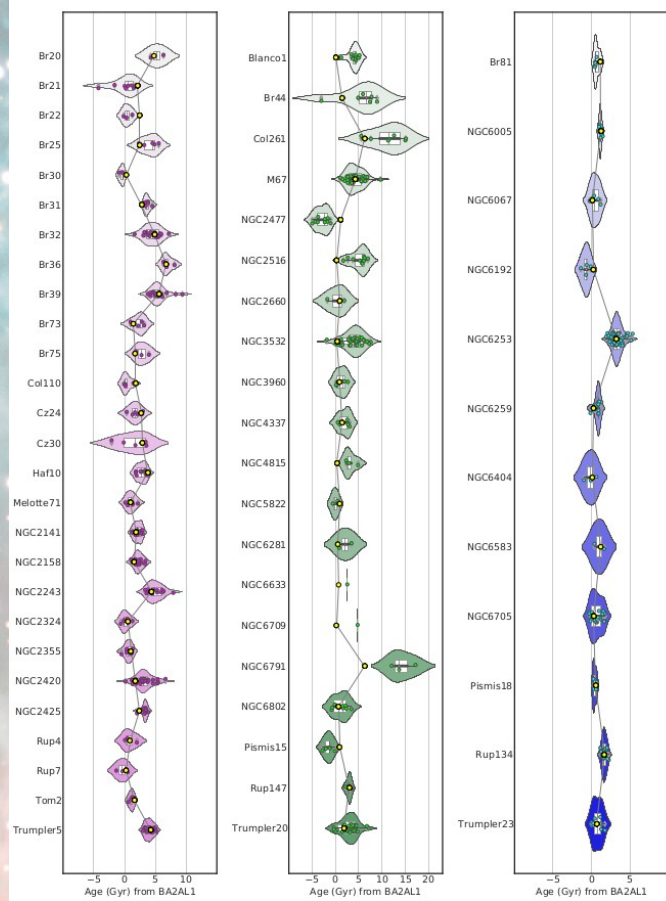


Viscasillas,
Magrini+22 (A&A)

Test 1: recovering the ages of the open clusters



Test 2: recovering the ages of member stars

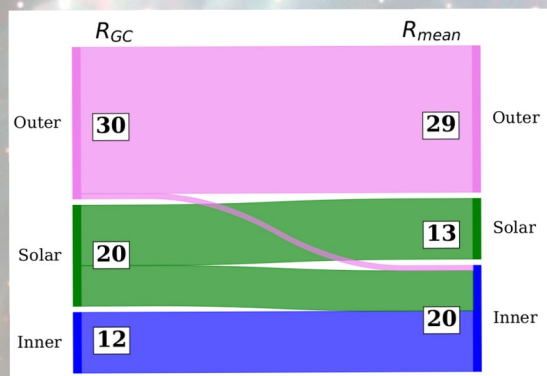
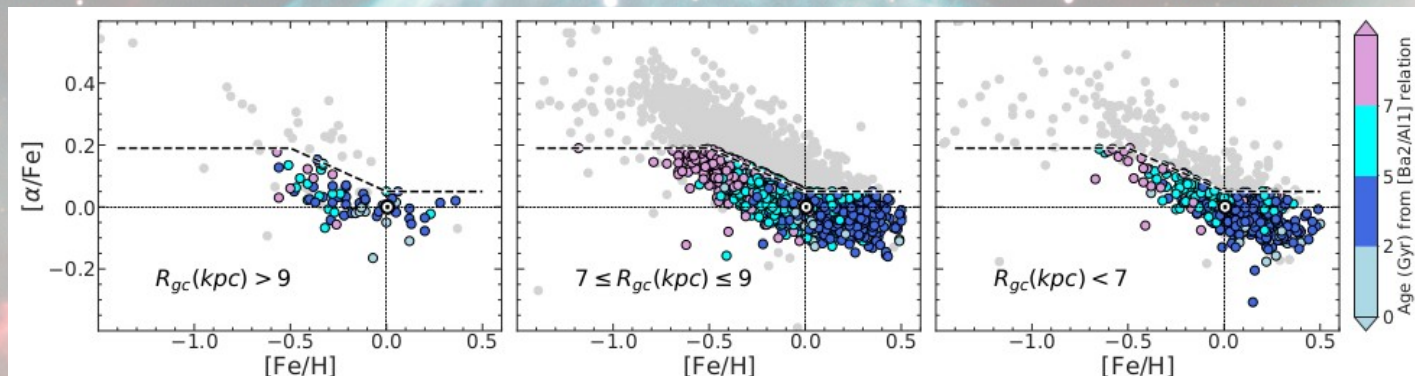


Viscasillas, Magrini+22 (A&A)

$$\text{Age} = m_1' \cdot [s/\alpha] + m_2' \cdot R_{gc} + m_3' \cdot [\text{Fe}/\text{H}] + c'$$

Test 3: infer ages of field stars

Viscasillas, Magrini+22 (A&A)



oldest stars: higher $[\alpha/\text{Fe}]$ ratio
youngest stars: solar or slightly sub-solar ratios

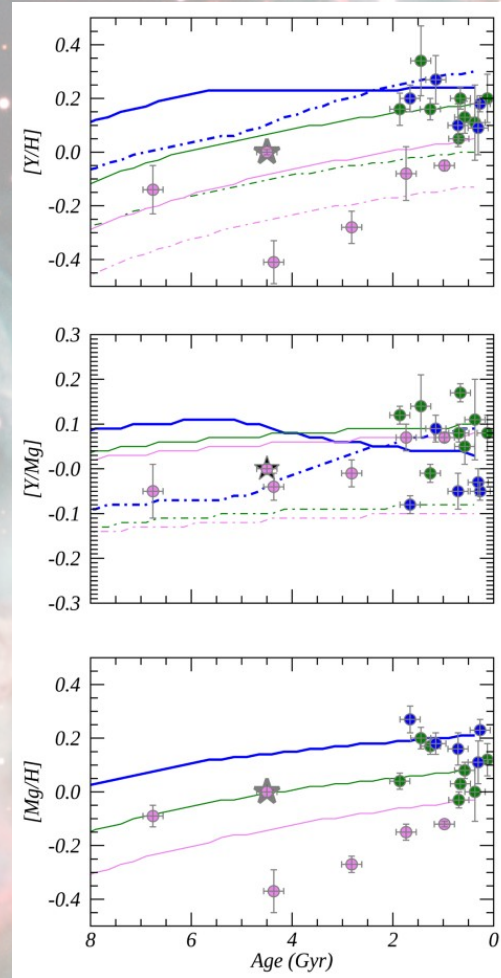
effects of migration

A theoretical explanation

Magnetic-buoyancy-induced mixing in AGB stars.

Less efficient production of Y at high metallicity, which affects this ratio on the inner disc.

MAGN yields (continuous)
FRUITY yields (dot-dashed)



Take home messages:

- Is possible to calculate ages directly from abundances using chemical clocks relations.
- The s-process elements Y and Ba are the most age-sensitive.
- There are no unique relations, they are different for each chemical clock and in each galactic region.
- The relations have a strong dependence on the galactocentric distance.
- $[\text{Ba}/\text{Al}]$ and in general abundance ratios involving barium, provide the best recovering factor.
- Able to reproduce the ages of the individual member stars with a precision better than 2 Gyr.
- Use the relations only on samples with the same characteristics (population, metallicity ranges, and galactocentric distance) and only consider ages in the range covered by the relationships.

Next steps:

The *r*-process, GES paper 115:

“The Gaia-ESO survey: placing constraints on the origin of
r-process elements”

Van der Swaelmen, Viscasillas Vázquez et al. (2022)
(A&A, recommended for publication)

To be continued...

at EAS SS15

“The Renaissance of Open Cluster Science with Gaia”

See you in Valencia!

Gracias

